





### Tamás Borkovits Baja Astronomical Observatory of Szeged University

(Not so) hierarchical stellar multiples with the eyes of *Kepler, TESS* et al.







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Importance of hierachical triples (and multiples)

### In the beginning...:

- Formation of close binaries (not only KCTF)
- Circumbinary planets
- Hot Jupiters?

### ...in the middle...:

- Sources of permanent perturbations (not only gravitational)
  - perfect sources of information

### ... at the end:

- Exotic terminal states of stellar evolution (e.g. binary pulsars
- And, even for not so exotic ones
  - E.g., what drives two white dwarfs to merge?





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Third body perturbations in a hierachical triple system

### Stellar three-body problem (Harrington, 1968, 1969)

Perturbations on three different time-scales

Classifications of periodic perturbations	Period	Relative amplitude	
Short period perturbations	~ P <sub>1</sub>	$\sim (P_1/P_2)^2$	
Long period perturbations	~ P <sub>2</sub>	$\sim P_1/P_2$	
"Apse-node" terms	$\sim P_{2}^{2}/P_{1}$	1	

Note: This is the classification introduced by Brown, 1936 for his Lunar-theory. Classification and nomenclature based on the planetary theory departs!



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Third body perturbations in a hierachical triple system

Vast majority of the triple system studies concentrate on the apse-node, or secular effects. Their importance and significance is out-of question.

### BUT what is the significance the shorter time-scale perturbations?

- theoretical: despite their low amplitude, they can substantially alter the secular dynamical evolutionary tracks of individual systems (Luo et al., 2016)

- practical (observational): it is, what we can observe and measure directly (thanks primarily to *Kepler* in the past, and to *TESS* in these days)

Discovery of more CHT-s Dynamical determination of masses & 3D configrations



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### **Eclipsing binaries in hierarchical triples – The Kepler revolution**

### <u>Two substantial giveaways -> an extended Royal Road</u>

**1.** Well-observable, very short time-scale perturbations





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For further details see: Borkovits et al. 2015, MNRAS, 448, 946 (analytic description) Borkovits et al. 2016, MNRAS, 455, 4131 (application for *Kepler* triples)



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The Royal Road(s): 1) Eclipse Timing Variation Analysis

In conclusion: information in dynamical ETV

- LTTE + dynamical perturbations:
  - light-travel time:  $P_2^{}$ ,  $a_{_{AB}}^{}$ sini $_2^{}$  [in km],  $e_2^{}$ ,  $\omega_2^{}$ , f(m<sub>c</sub>)
  - third-body perturbations:



- (P<sub>2</sub> time-scale): P<sub>2</sub>, m<sub>c</sub>/M, e<sub>2</sub>, (
$$\Omega_1 - \Omega_2$$
), i<sub>mut</sub>, i<sub>0</sub>, g<sub>2</sub>  
e<sub>1</sub>,  $\omega_1$ , g<sub>1</sub>, h, j<sub>1</sub>, j<sub>2</sub>

-  $(P_2^2/P_1 \text{ time-scale}): e_1, \omega_1, e_2, u_{mut}, g_1, g_2, h, m_c/M$ 

(apsidal motion, orbital plane precession)

The yellow quantities were <u>almost</u>, while the purple ones were <u>completely</u> unknown for compact triples before *Kepler*-era, although they are very important for dynamical evolution studies



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SCIENTIARUM SHIPPING

The Royal Road(s): 1) Eclipse Timing Variation Analysis

### Some statistics for 222 triples (including 160 pure LTTE systems, too)

- Shortest outer period systems
  - The tightest binaries (mostly overcontact systems) have no very close outer companion(s)? (The brown region is almost empty.)

Some implication for their formation mechanism?





Short time-scale perturbations and their modelling in compact hierarchical triple systems



**The Royal Road(s): 1) Eclipse Timing Variation Analysis** 

### Some statistics for 62 triples

• Mutual inclination:

(Only for systems with dynamical perturbations.)

- (Second) peak at  $i_m \sim 40^\circ$  (as Fabrycky & Tremaine, 2007 predicts), but the inner period distribution is inconsistent with their predictions





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#### **Eclipsing binaries in hierarchical triples – The Kepler revolution**

### <u>Two substantial giveaways -> an extended Royal Road</u>

2. Extra eclipses – Trinity-like systems (HD 1810681)



For further details see: Derekas et al. 2011, Science, 332, 216 (discovery paper) Borkovits et al. 2013, MNRAS, 428, 1656 (lc+RV+ETV analysis) Fuller et al. 2013, MNRAS, 429, 2425 (tidally forced oscillations)



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Extra Eclipser's Hall of Fame\*

11 - Kepler prime mission\*\*
2 - CoRoT
2 - K2

\* As of 15. October, 2018\*\* Not counting Kepler`s circumbinary planets



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- **11** *Kepler* prime mission\*\*
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- 1 AAVSO observers

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The Royal Road(s): 2) (Spectro –) Photodynamical Analysis of ...

... (mostly, but not exclusively) multiply eclipsing multiple systems

### **Photo-dynamical:**

#### **Classical** eclipsing binary lightcurve analysis:

- stellar motion purely Keplerian and described mostly only in 2D!!! (fix orbital plane)
- only exception: apsidal motion  $\rightarrow$  linear in time
  - (... khm... in the new versions of the WD code a Keplerian LTTE is also built in)
- presence of a third star can be accounted only with the use of an additive quantity of third light (wavelength-dependent)
- no outer eclipses

Eclipsing binaries in compact hierarchical multiples or, planetary systems:

- need a complete, 3D <u>dynamical</u> treatment of the motion,

 i. e. numeric integration of the three, four, etc.-body motion, even including relativistic and/or tidal forces (and even integration of the stellar spin equations)
 – handling of different kinds of extra eclipses, etc.

Some examples: Carter et al. 2011, Science, 331, 562 (KOI-126); Welsh et al. 2012, Nature, 481, 475 (K-34,-35); Dawson et al. 2014, ApJ 791, 89 (K-419);Orosz 2015, ASPC 496, 55 (KIC 07668648, 10319590), etc.



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The software package: Lightcurvefactory

### More precise physical models:

Radiative and atmospheric models – identical with WD, PHOEBE

 → to be extended in the near future

 Inclusion of oscillations, tidal effects – preliminary models – 2011-/2018 –

 Inclusion of some relativistic effects (e.g. Doppler-boasting) – 2011 –

### Extended geometry and dynamics...

Three-, and even four-body eclipses – 2011/2017 – Inclusion of three- and four-body dynamics (with numerical integration) – 2017 (inclusion of further hierarchic, and planetary configurations)

### Simultaneous analysis of different kinds of data:

Multi-band photometry - 2014 -RV- and ETV-curves - 2017 -SED+isochrones - last week



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**The Royal Road(s): 2) Photodynamical Analysis of Triply Eclipsing Systems...** 

#### A recent example: EPIC 249432662



A ~15-mag star in Scorpius, observed during Campaign 15 of the K2 mission

~3% deep eclipses reveal a slightly eccentric ~8.2-day period EB

An almost 50% deep, more than 2-day-long, extra fading with complex structure

ETV analysis reveals rapid period variations



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The Royal Road(s): 2) Photodynamical Analysis of Triply Eclipsing Systems...

**Spectroscopic follow-up** (KECK I HIRES + McDonald Observatory)



Only one, probably MS K star is seen in the spectra

Joint photodynamical analysis of K2 lightcurve + ETV curve + RV data:

The fourth RV point obtained at 22 March 2018 allowed us to constrain the outer period with ~2-3-day accuracy.





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The forthcoming outer eclipses should be occurring within 5 – 10 days of that time!



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# SCIENTIAROM STREET

### The Royal Road(s): 2) Photodynamical Analysis of Triply Eclipsing Systems...





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The Royal Road(s): 2) Photodynamical Analysis of Triply Eclipsing Systems...



Most interesting results: - extremely flat system:  $i_{mut} = 0.189^{\circ} \pm 0.124^{\circ}$ - similar mass-ratios:  $q_{in} = 0.876 \pm 0.012$ ;  $q_{out} = 0.877 \pm 0.044$ 

Further details in Borkovits, Rappaport, Kaye et al., 2019, MNRAS 483, 1934

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The Royal Road(s): 2) ... and quadruple system (both 2+1+1 & 2+2)

### An example for the usefulness of the desk drawer observations : EPIC 210966582



A relatively bright (V ~ 10.84), red dwarf in the solar neighbourhood, located in the auspicious region of Cancer which was observed in three K2 campaigns (C05,16,18). K2 observations show cyclic ETV with a period of ~59<sup>d</sup>. Slight eclipse depth variations are also visible...



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### An example for the usefulness of the desk drawer observations : HIP 41431



Independent spectroscopic observations by Julius Sperauskas & Andrei Tokovinin, with different instruments (mainly VUES & CHIRON): an SB 3 system formed by three very similar stars



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Andrei was looking for former RV measurements, and he realized that Dave Latham at CfA had been observed this system continuously for two decades!

– According to these old RV data, this system is indeed a 2+1+1 quadruple system, with an outmost period of ~  $P_3$ =1436<sup>d</sup>.



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EPIC 212096658ABC RV curve



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- According to these old RV data, this system is indeed a 2+1+1 quadruple system, with an outmost period of ~  $P_3$ =1436<sup>d</sup>.
- It explains also the systematic residual of the ETV in the three-body photodynamical fit.



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Primary (left) and secondary eclipses (right) observed this year with the 50-cm RC telescope of Baja Astronomical Observatory: eclipse depth variations are evident, and in nice accordance with the predictions of the 2+1+1 spectro-photodynamical model.

Further details in Borkovits, Sperauskas, Tokovinin et al., 2019, MNRAS 487, 4631



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#### The future is in progress...





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## Thank you for the attention!