

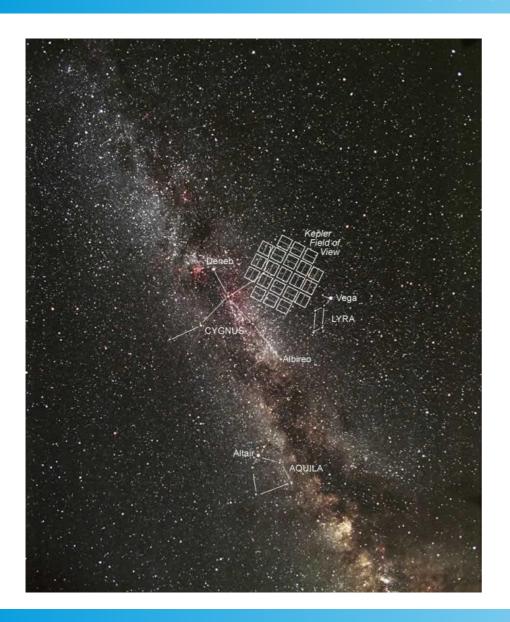


Eclipsing binaries hiding in the background: The Kepler Pixel Project

Adrienn Forró¹²³, Róbert Szabó¹², Tamás Hajdu¹²³

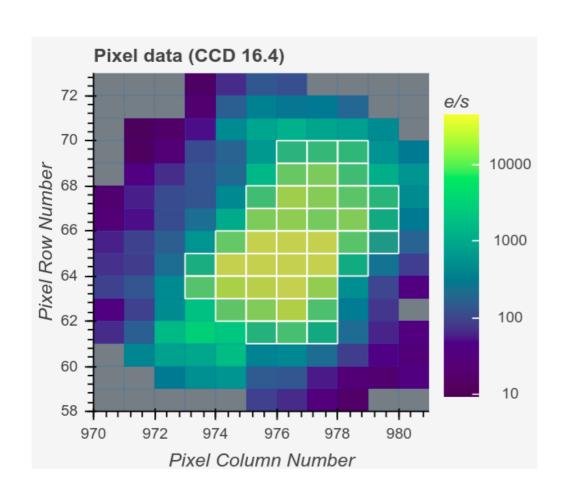
- 1 Konkoly Thege Miklós Astronomical Institute, Research Centre for Astronomy and Earth Sciences, Hungarian Academy of Sciences
- 2 MTA CSFK Lendület Near-Field Cosmology Research Group
- 3 Eötvös Loránd University, Department of Astronomy

Introduction



- Kepler's original mission: 2009-2013
- continuous observations for ~150000 stars

The Kepler Pixel Project

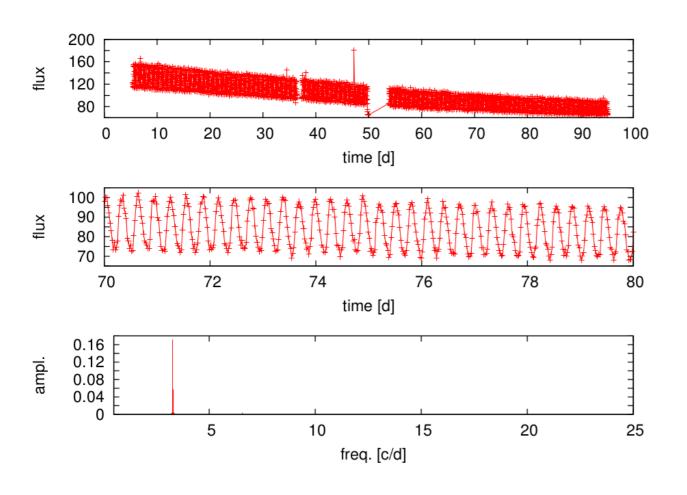


Q4: 2009 Dec – 2010 March

Primary goal:

to find new variable stars in the background pixels

The Kepler Pixel Project

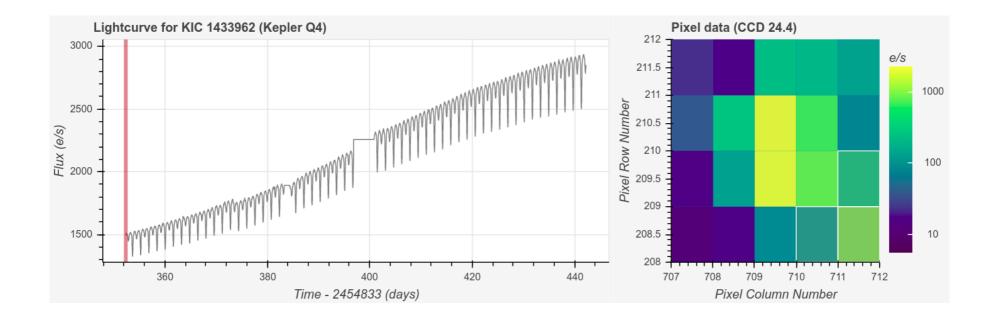


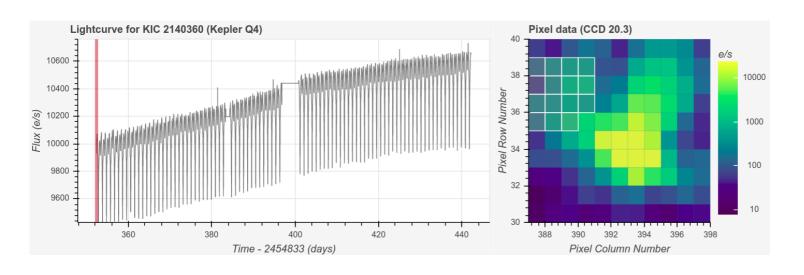
Searching criteria:

- period: 0.25 1 day
- the presence of at least 2 harmonics

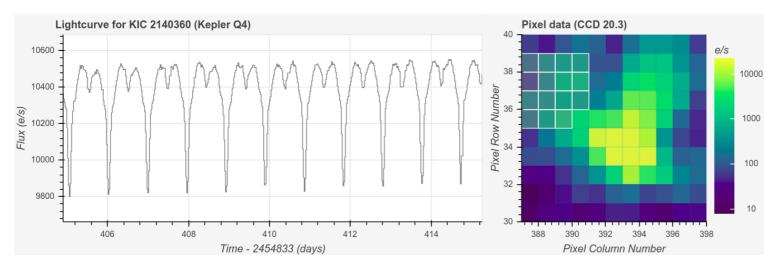
- ~ 12500 candidate pixels in total
- mostly eclipsing binaries, 1272 TPFs

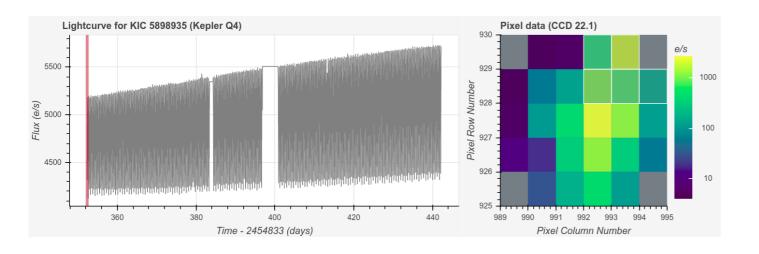
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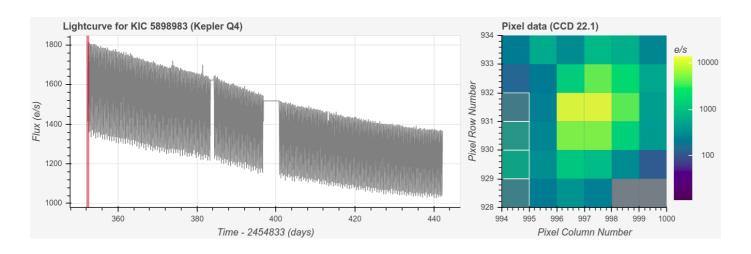


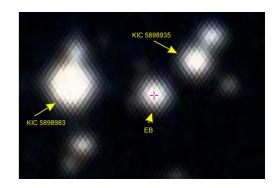






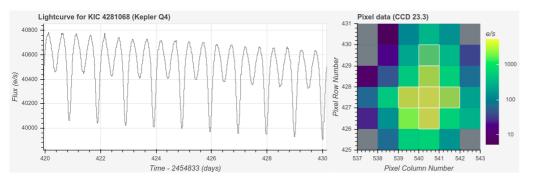


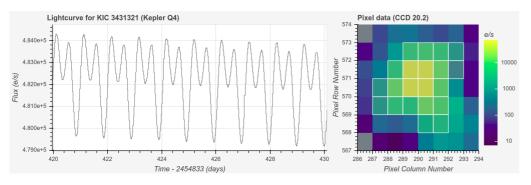




P=0.5787d

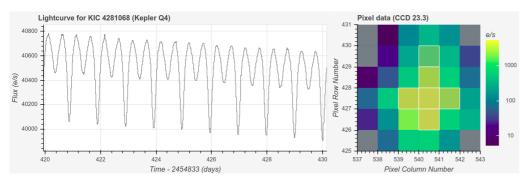
- we looked for duplicates/contamination 1 candidate
- we cross-matched them with the Kepler Eclipsing Binary Catalog (by the Villanova University) 495 candidates

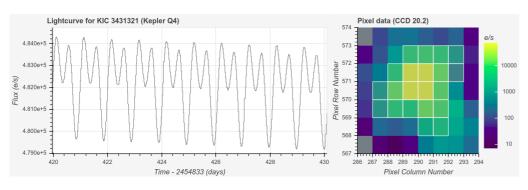




P=1.015d

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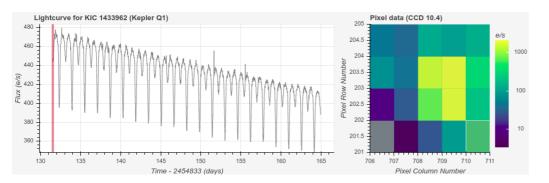




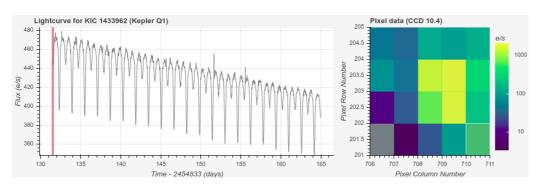
- 776 new EB candidates in the original Kepler field
- substantial increase in the number of binaries, the catalog lists 2922 EBs

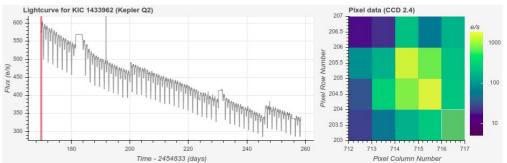
P=1.015d

- some quarters might be missing
- TPFs from different quarters are not necessarily the same size

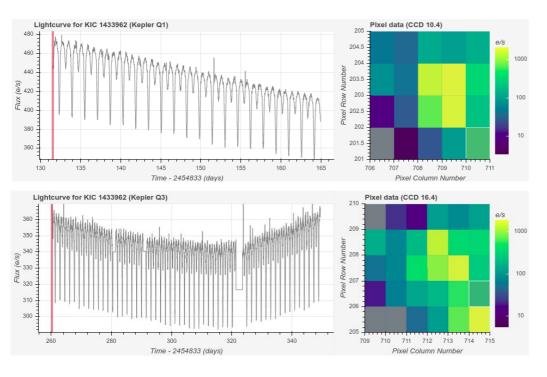


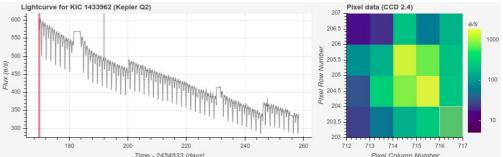
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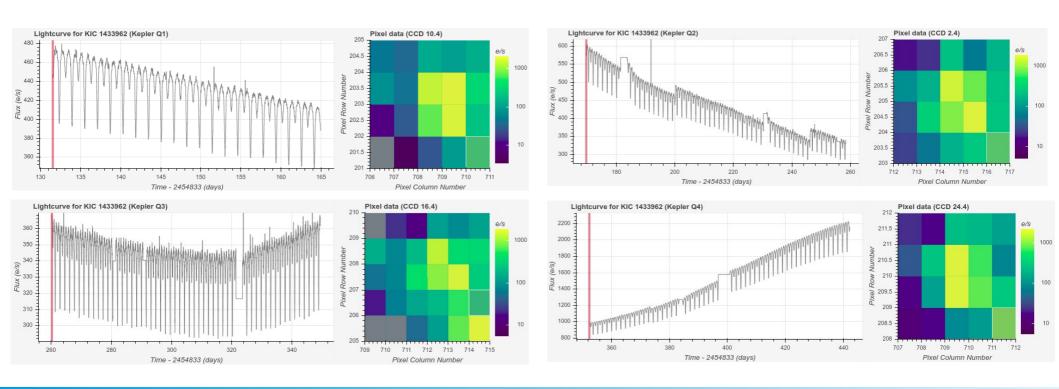


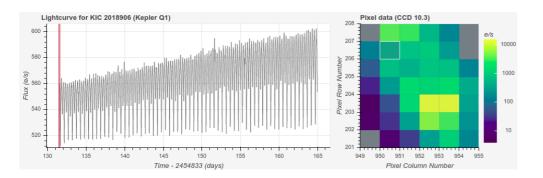
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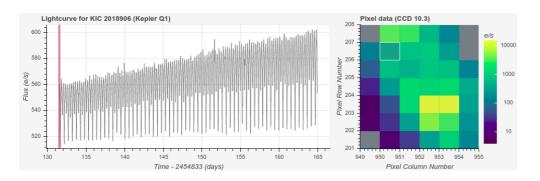


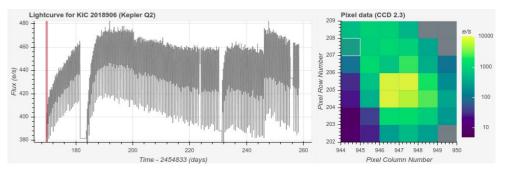


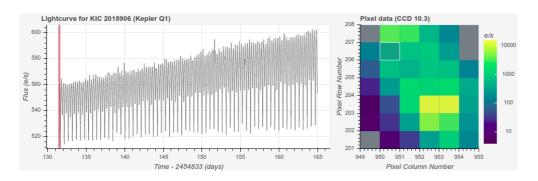
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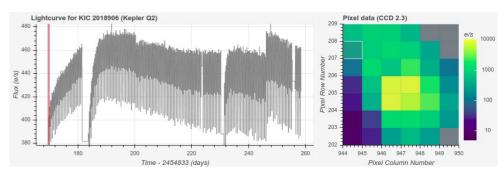


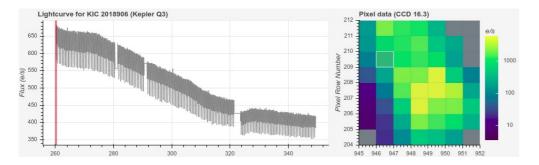


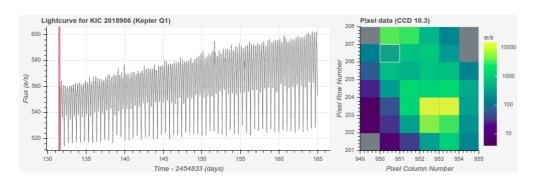


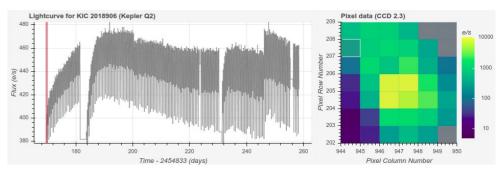


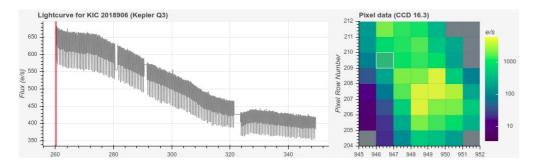


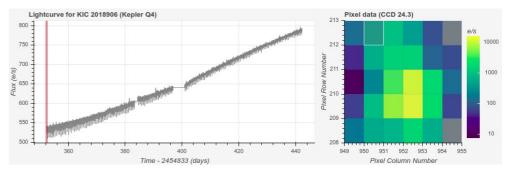


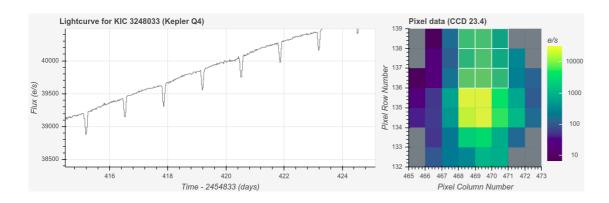


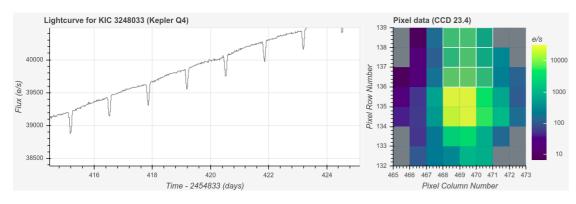




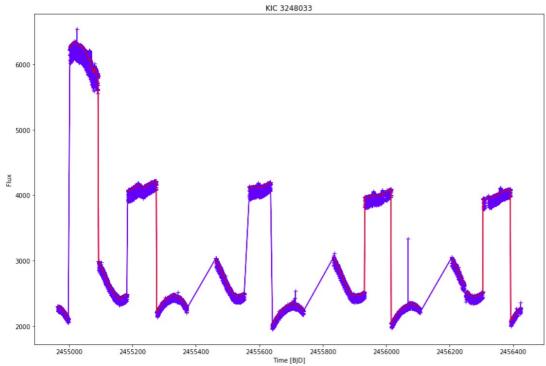


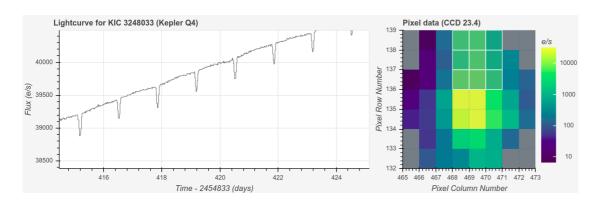




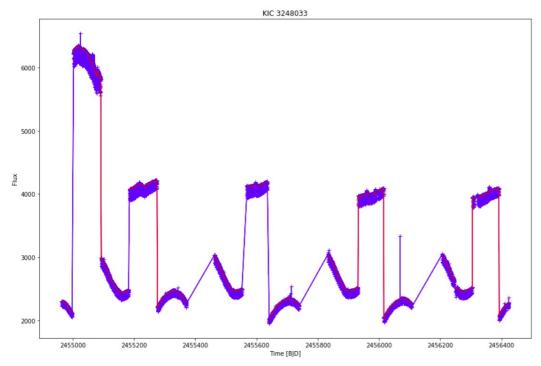


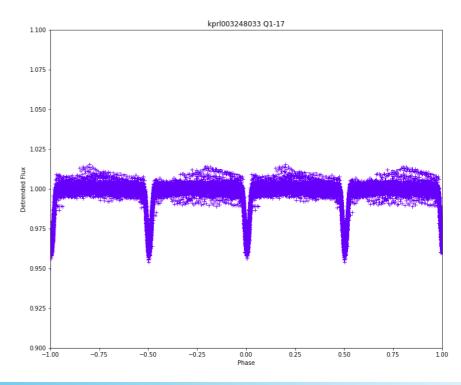
Detrending with Wötan (Hippke et al. 2019)

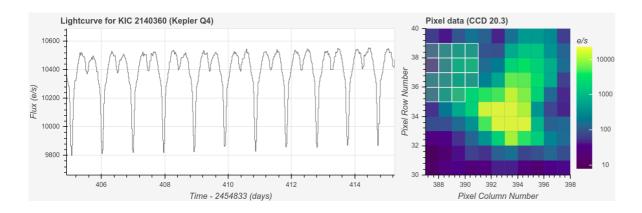


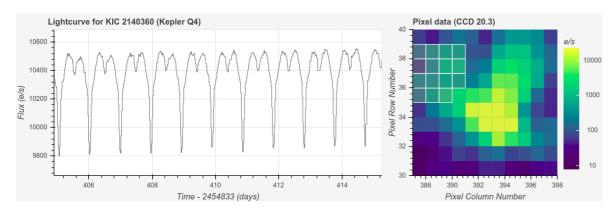


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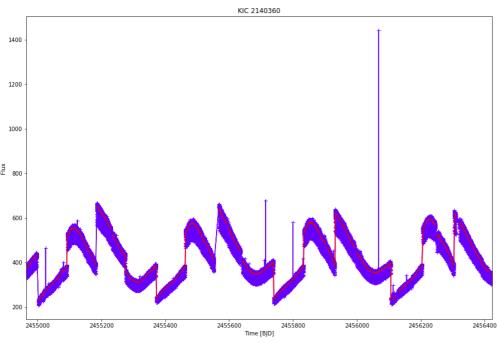


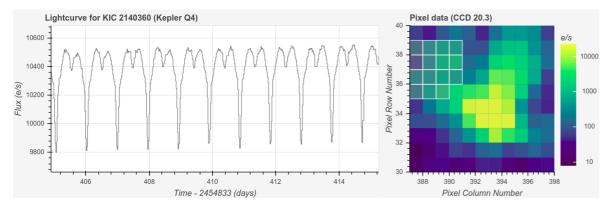




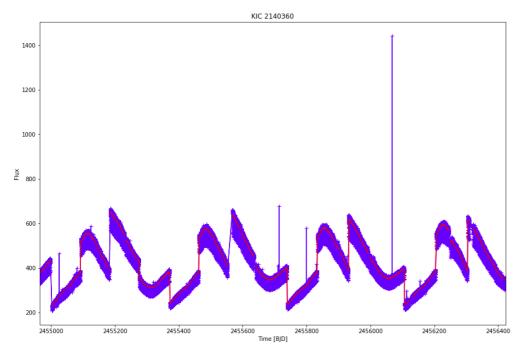


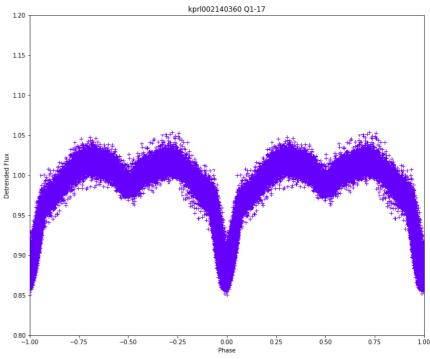
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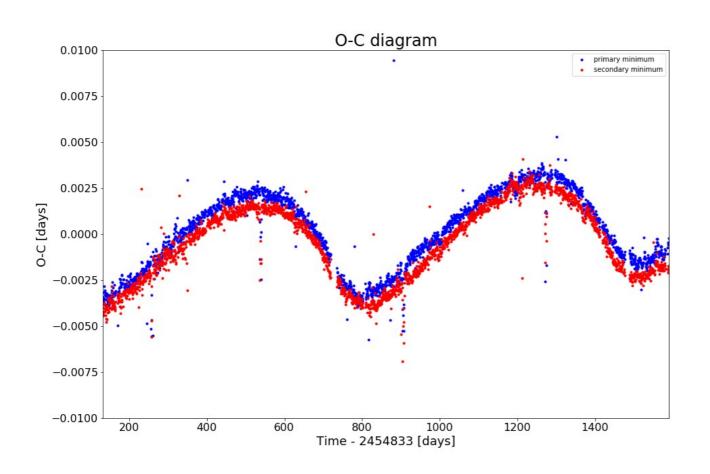


Detrending with Wötan (Hippke et al. 2019)





Triple-star system candidates?



Candidate located on 2 TPFs: KIC 5898935 & KIC 5898983

- we were searching for variable stars in the background: P=0.25-1d at least 2 harmonics

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- search in other period ranges as well: P<0.25: John Bienias